

The LSO for our friend Manolo Collados Vera

the LSO group

2022/12/03, LSO, Slovakia



LSO ZEISS (Lyot) coronagraph

















Mount

ZEISS mount VII a



ZEISS mount VII a

- Theory ~ reality: all parts are bended depending on relative orientation of mechanical structures to the gravity vector direction → differential bending of the coronagraph → the solar disk image shift in the coronagraph focal plane
- Pointing systems correcting also for such effects are needed
- LSO: no active and/or adaptive optics



Pointing

AISAS pointing for single coronagraph



AISAS pointing for single coronagraph













- NCAR/HAO product (S.Tomczyk) for the EU structural funds for science in Slovakia
- Tunable 4-stage Lyot filter + polarimeter
- Pre-filters + calibration optics
- Dichroich mirror
- Polarizing splitting cubes
- Pairs of detectors for the VIS and the IR
- Optics, heating, cooling

- The Lyot filter:
 - birefringence: calcite, stages from two calcite partsfr
 - broadband polarizers: VIS700BC4 by CODIXX
 - super-achromatic $\lambda/2$ plates: APSAW by ASTROPRIBOR (0.7-1.5 λ_{o})
 - FWHM: 0.03-0.13 nm (530-1083 nm)



- The polarimeter:
 - polarization modulator: scheme from HAO Prominence Magnetometer (ProMag)
 - 2 ferro-electric liquid crystals (FLC) + fixed retarder not achromatic → value of retardation and orientation are selected with a Monte Carlo procedure which optimizes the Stokes modulation efficiency over a very broad



- the polarimeter calibration:
 - HAO has proposed is based on the following articles:
 - del Toro Iniesta and Collados, Appied Optics, 39, 1637 (2000)
 - Tomczyk, Casini, de Wijn, Nelson, Applied Optics, 49, 2010
 - two HAO internal documents of Tomczyk and De Wijn
 - HAO has provided some IDL routines (but with problems reported)
 - AISAS:
 - the proposed polarimeter calibration has not been applied yet in plan for 2023
 - The polarimetric calibration data are regularly acquired

CoMP-S: filter module



CoMP-S: camera module



CoMP-S: camera module



CoMP-S: the spectral lines

- Coronal lines: Fe X 637 nm (Fe XIV 530 nm, FE XIII 1074+1079 nm)
- Prominence lines: He D3 587 nm, He I 656 nm, Ca II 854 nm (He I 1083 nm)



CoMP-S: electronics



CoMP-S: electronics + computers



CoMP-S: operation



CoMP-S: operation



CoMP-S: operation



CoMP-S: example frames

-2664 1080-1000-900-854 nm (100 ms) 800-700-600-500-H=7123 1080-1000-656 nm (50 ms) 900-800-700-600-500-H-6917 1080-1000-587 nm (500ms) 900-800-700-600-500-

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-2250 -2000

-1750

-1500 -1250

-1000

-750 500

-6000

-5000

-4000

-3000

-2000 1000

-6000

-5000

4000

-3000

-2000

1000

CoMP-S: example phot_red H alpha



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CoMP-S: example H alpha scan



- Observing programs:
 - old: H alpha tornadoes, Masses of proms
 - actual: He D3 in quiet/eruptive proms
 - future: 2023 open call
- Projects under development/improvements: CoMP-S instrument, SCD instrument, pointing system, LSO pipeline
- New projects for time to come: LSO \rightarrow VSO, SLED instrument, dome motion, dome a la "THEMIS",...



LSO group and work

LSO group and work

- LSO group:
 - handyman
 - observing assistants: 2 astronomers and 2 engineers
 - support from the headquarters: workshop, sys admin
- LSO duty shifts (~8 days):
 - cablecar transportation including all your things
 - cooking
 - observing or testing of instrumentation
 - own instrumental or scientific projects
 - day and night security officer as well
 - person on duty



LSO for students

LSO (summer or autumn) internships

- Student interested in a stay at the LSO to learn more about the g-b coronagraphic observations are welcome! Just send an email...
- Conditions: stay at the LSO is free of charge, travel including the cable car ticket and foodstuff is up to you.
- Small groups are preferred



2022 student Mark Morris @ LSO

In fine

